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CAUSAL ADS - SPACETIMES

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Classical and quantum gravity
in 3 dimensions.

AdS: anti-de Sitter space ($n=3$)

time oriented, oriented $G = SO_0(2,2)$

$M^{(2n)}$ locally AdS (Einstein in the void, $\Lambda < 0$)

Example: $M = \Gamma \backslash \Omega$ $\Omega \subset \widetilde{\text{AdS}}$
 $\Gamma \subset G$

Our concerns:

① Given $\Gamma \subset G$ discrete, find $\Omega \subset \widetilde{\text{AdS}}$
such that $\Gamma \backslash \Omega$ manifold

② Description of BTZ (multi-) black-holes
 $\Leftrightarrow (\mathbb{S}^2, PSL_2\mathbb{C})$ -structure \Leftrightarrow Ku-Pi measured laminations

For ①: Comparison with \mathbb{H}^3 :

If $\Gamma \subset SO_0(1,3) \simeq \text{Isom}(\mathbb{H}^3)$:

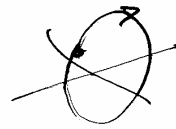
$\Gamma \backslash \mathbb{H}^3$ (complete) \mathbb{H}^3 -manifold

+ limit set $\Lambda_\Gamma \subset \partial\mathbb{H}^3 = \mathbb{S}^2$

for $\Gamma \subset G = \text{Isom}(\text{AdS})$: NO!

CAUSALITY NOTIONS

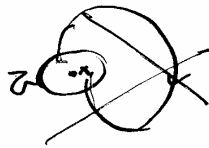
- CAUSAL SPACETIMES non-spacelike closed curves forbidden:

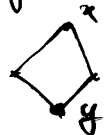


- STRONGLY CAUSAL SPACETIMES every point x admits a "causally convex" neighborhood \mathcal{U}



Alexandrov Topology
manifold Topology



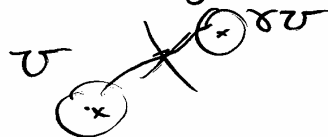
- Globally hyperbolic spacetimes: (maximality) \exists Cauchy hypersurface \Leftrightarrow compactness of 

Γ acting on strongly causal spacetime:

- CAUSAL ACTION: for every x , every δ :



- STRONGLY CAUSAL ACTION: every point x admits a neighborhood \mathcal{U}



MEANINGFUL FOR EVERY PSEUDO-RIETANNIAN METRIC.

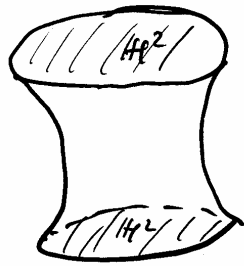
- ALL RIETANNIAN MANIFOLDS ARE GLOBALLY HYPERBOLIC
- FOR RIETANNIAN ACTIONS: causal action \Leftrightarrow free
strongly causal action \Leftrightarrow proper

anti-de Sitter AdS

$$Q = -u^2 - v^2 + x_1^2 + x_2^2 \text{ on } \mathbb{R}^{2,2}$$

$$\text{AdS} = \{Q = -1\}$$

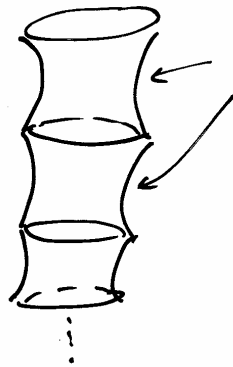
projects on ADS in \mathbb{RP}^3 (Klein model)



hyperboloid

AdS is NOT causal

$\tilde{\text{AdS}}$



"affine domains"

$\tilde{\text{AdS}}$ is STRONGLY CAUSAL

(but not globally hyperbolic)

$$\tilde{\text{AdS}} \hookrightarrow \tilde{\text{Ein}}_3 \approx \mathbb{S}^2 \times \mathbb{R} \quad ds^2 - dt^2$$

Einstein Universe
↑
lightlike geodesics

CONFORMAL EMBEDDING
i.e. respect causality relation.

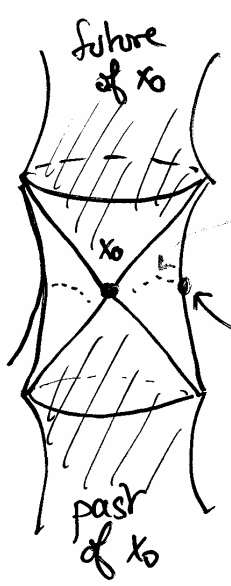
$$\partial \tilde{\text{AdS}} \approx \hat{\text{Ein}}_2 \subset \tilde{\text{Ein}}_3 \quad \text{"CONFORMAL BOUNDARY"}$$

(compare: $\partial \text{H}^3 \approx \mathbb{S}^2 \subset \mathbb{S}^3$)

$\tilde{\text{Ein}}_n$ IS GLOBALLY HYPERBOLIC.

$$\Gamma \subset \text{Isom}(\widehat{\text{AdS}}) \approx \widetilde{SO}_0(2,2)$$

If $x_0 \neq y_0$ in $\widehat{\text{AdS}}$ $\forall \gamma \in \Gamma$
 Γ non causally related

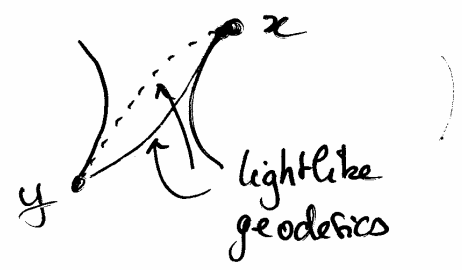


orbit $\Gamma \cdot x_0$

$$\Lambda_\Gamma = \overline{\Gamma \cdot x_0} \cap \partial \widehat{\text{AdS}}$$

DEFINITION: Γ admissible if Γ preserves some generic achronal closed subset Λ of $\widehat{\text{Ein}}_2$

(Generic: does not contain



ELEMENTARY CASES:

- $\Lambda = \emptyset, 1$ or 2 points
- $\Lambda = /$ (lightlike segment) extreme
- $\Lambda = \wedge$ (2 lightlike segments) conical

$$\text{Isom}(\text{ADS}) \simeq \text{SO}_0(2,2) \simeq \text{PSL}(2, \mathbb{R}) \times \text{PSL}_2(\mathbb{R})$$

$$\begin{array}{ccc}
 & \nearrow e_L & \text{PSL}_2 \simeq \text{PSL}_2 \times \{ \text{id} \} \\
 \Gamma & \hookrightarrow & \text{Isom}(\text{ADS}) = \text{PSL}_2 \times \text{PSL}_2 \\
 & \searrow e_R & \text{PSL}_2 \simeq \{ \text{id} \} \times \text{PSL}_2
 \end{array}$$

(Γ non abelian) Admissible

\Uparrow
 e_L, e_R faithful, discrete, semi-conjugate

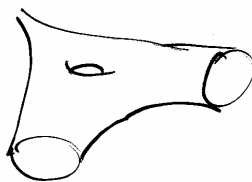
$$\Sigma_L = \mathcal{C}_L(\Gamma) \backslash \mathbb{H}^2 \quad \text{homeo.} \simeq$$

$$\Sigma_R = \mathcal{C}_R(\Gamma) \backslash \mathbb{H}^2$$

\swarrow means



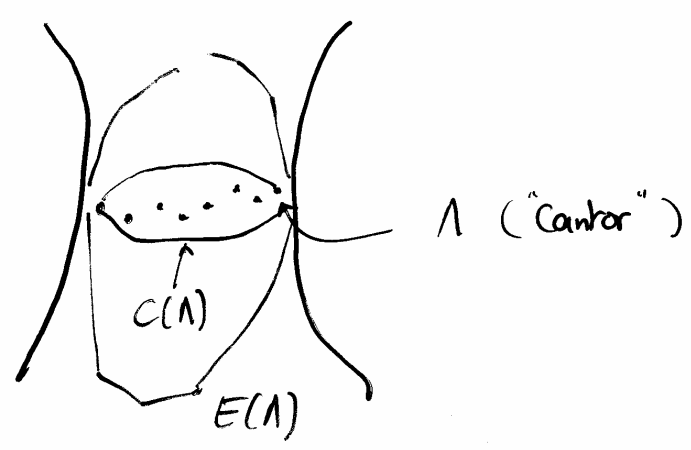
(marked surface)



(marked surface)

$$(\Sigma_L, \Sigma_R) \in \underline{\text{Teich}} \times \underline{\text{Teich}}$$

$\Lambda \subset \widehat{Ein}_2$ closed achronal generic $\Rightarrow \Lambda \subset$ affine domain



- $C(\Lambda)$: convex hull of Λ
- $E(\Lambda)$: Invisible domain from Λ in AdS
 $(E(\Lambda) = \{x \in \text{AdS} / \forall p \in \Lambda, x \not\ll p\})$
- $\Omega(\Lambda)$: Invisible domain from Λ in Ein_2
- $\bar{E}(\Lambda) = E(\Lambda) \cup \Omega(\Lambda)$

THEOREM: Λ generic achronal non-elementary subset preserved by Γ .
 THEN, THE ACTION OF Γ ON $\bar{E}(\Lambda)$ IS PROPERLY-DISCONTINUOUS AND STRONGLY CAUSAL

(Γ torsion-free \Leftrightarrow free action)

$$M_\Lambda(\Gamma) = \Gamma \backslash E(\Lambda)$$

$$\bar{M}_\Lambda(\Gamma) = \Gamma \backslash \bar{E}(\Lambda)$$

Λ ELEMENTARY? CASE-BY-CASE ARGUMENT,
NOT ALWAYS TRUE. IMPLY Γ abelian

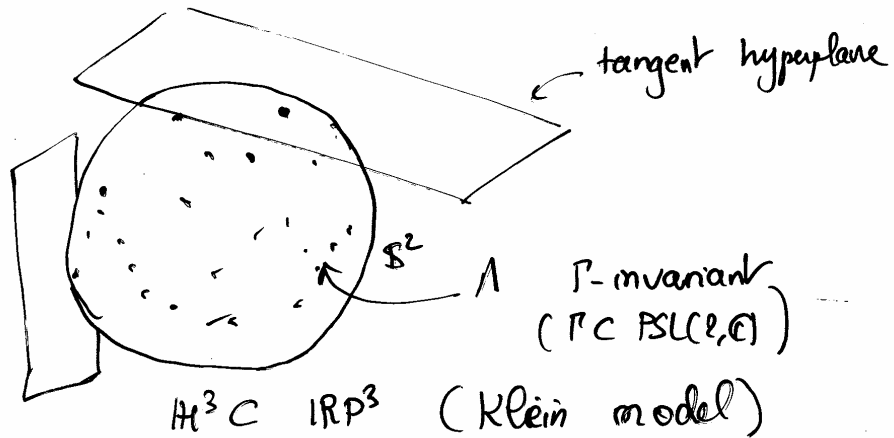
ANYWAY:

IF Γ IS NON-ABELIAN, THEN

$$\Lambda_\Gamma = \{ \text{attractive fixed points} \}$$

THE MINIMAL Γ -INVARIANT ACHRONAL
SUBSET

COMPARISON WITH RIEMANNIAN CONTEXT:



$\mathbb{RP}^3 \setminus \overline{\mathbb{H}^3} \approx \text{de Sitter space(time)}$
 $\Rightarrow E(\Gamma) = \mathbb{H}^3 \cup \underbrace{\Omega(\Gamma)}_{\mathbb{S}^2/\Gamma} \cup E(\Gamma)$
 $E(\Gamma) \subset dS_3$
 $\Gamma \backslash E(\Gamma)$ Globally hyperbolic dS -space-time
 (Scannell)

AdS - case :

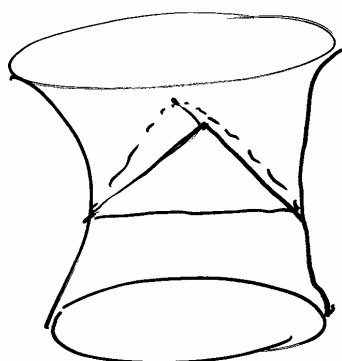
7.

Λ topological circle $\Leftrightarrow \Omega(\Lambda) = \emptyset$



$M_\Lambda(\mathbb{P})$ GLOBALLY HYPERBOLIC

Example Torus Universes:

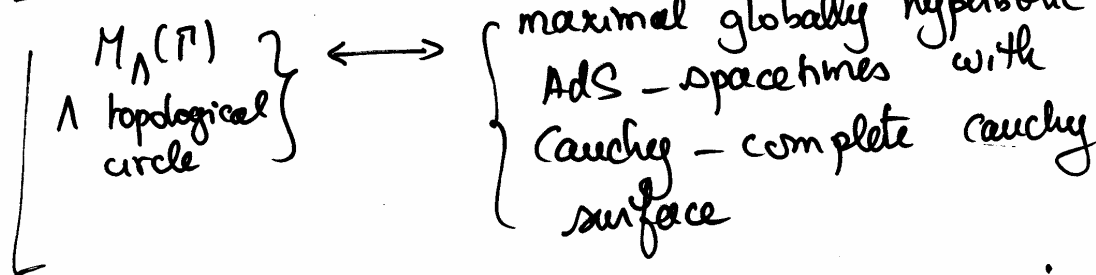


$C(\Lambda) = E(\Lambda) = \text{Tetraedron}$

\mathbb{P} abelian $\subset \mathbb{Z}^2$

$\mathbb{P} = \mathbb{Z}^2$: torus universe

THM (Bonsante - Benedetti)
(Barbot)



① \leftarrow : Mex : $\Sigma \subset \text{AdS} \Rightarrow \Sigma \subset \mathbb{H}^2$ distance increasing

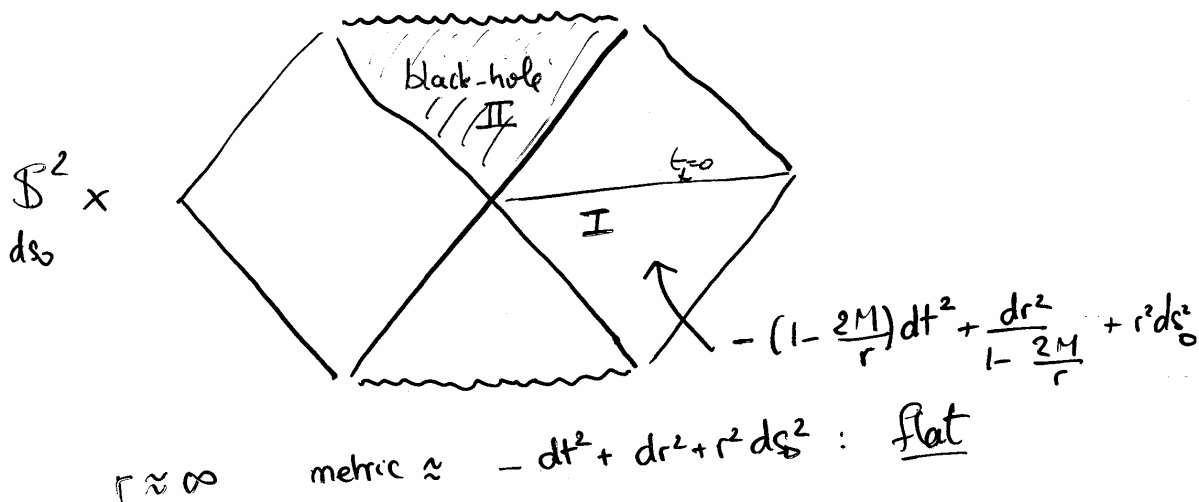
② \rightarrow : $M_\Lambda(\mathbb{P})$ contains a Cauchy surface?

(My proof: $\Sigma \subset \text{AdS spacelike} \Rightarrow \Sigma \subset \mathbb{H}^2 \times \mathbb{H}^2$)

Σ convex $\Rightarrow \Sigma \subset \mathbb{H}^2 \times \mathbb{H}^2$ distance decreasing

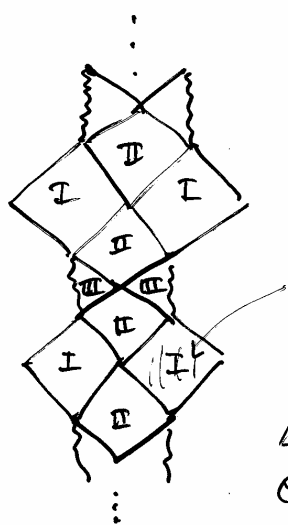
BLACK HOLES:

A quick insight into Schwarzschild black-hole
Kerr



OBSERVERS AT CONFORMAL BOUNDARY
VERY FAR

Kerr:



$$-(1 - \frac{2Mr}{\rho^2})dt^2 + \sin^2\theta \frac{4Mra}{\rho^2} dt d\phi + \rho^2 d\theta^2 + (r^2 + a^2 + \sin^2\theta \frac{2Mra}{\rho^2}) \sin^2\theta d\phi^2 + \frac{\rho^2}{\Delta^2} dr^2$$

$$\Delta = r^2 - 2Mr + a^2$$

$$\rho^2 = r^2 + a^2 \cos^2\theta$$

M : mass

a : angular momentum per unit mass

CONFORMAL BOUNDARIES

$$M \text{ Lorentzian} \Rightarrow \bar{M} = M \cup \underbrace{\partial M}_{\text{conformal boundary}}$$

"Universal":

$$\bar{M}_{\text{univ}} \text{ if for every } \bar{M} = M \cup \partial M$$

$$M \hookrightarrow \bar{M} \text{ extends to injection } \bar{M} \hookrightarrow \bar{M}_{\text{univ}}$$

IN GENERAL ?

FOR AdS-spacetimes:

$$E \subset \text{AdS}, \text{ define } \partial E = \text{Int}(\bar{E} \cap \text{Ein}_2)$$

$$\text{and } \bar{E} = E \cup \partial E$$

(AdS, Ein₂, Iso(AdS))-structure

THM: Any AdS-spacetime admits

└ a universal conformal boundary

(complete case: Frances)

⇒ can define observers, black-holes

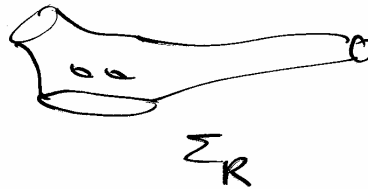
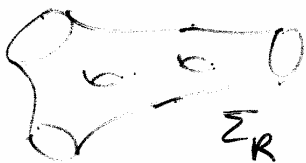
BTZ - multi black-holes

Bañados
Teitelbaum
Zanelli
+ Ammonberg, Brill,
Peldán, etc... 10.

$$\Gamma \subset SO_0(2,2) \text{ admissible}$$

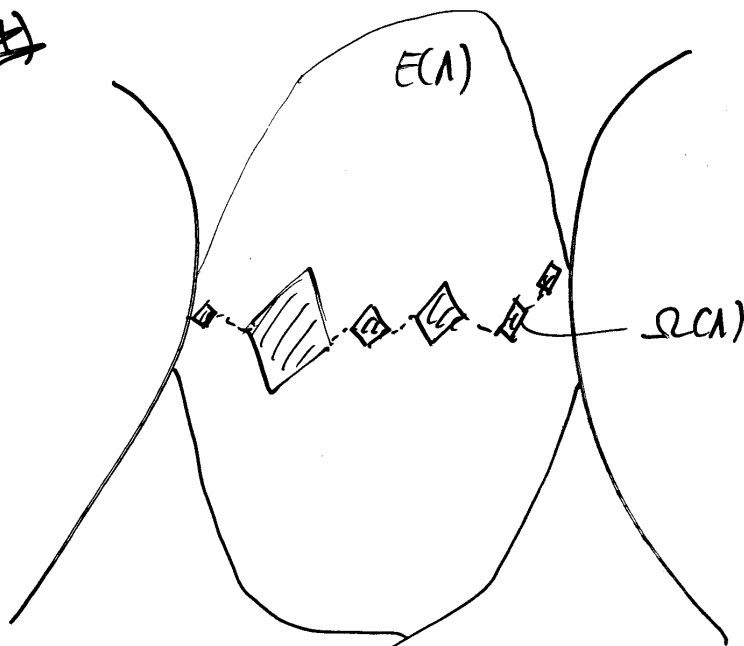
$$\updownarrow$$

$$\rho \in \text{Rep}^*(\Gamma, \text{PSL}_2)$$



$\Lambda = \Lambda_p$ limit set

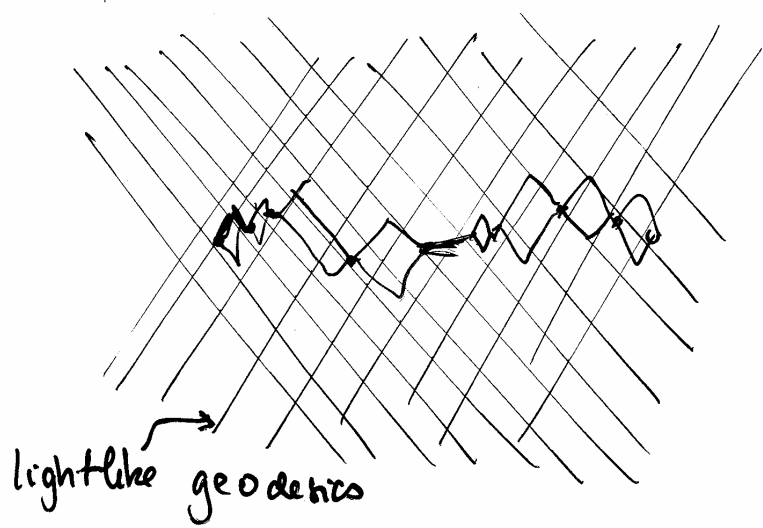
~~ECA~~



$$M_\Lambda(\Gamma) \subset \overline{M_\Lambda(\Gamma)}$$

$$\partial M_\Lambda(\Gamma) = \Omega CA = \text{observers}$$

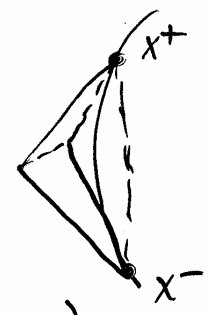
||



Λ^+
 Λ^- \subset Ein₂

$$E(\Lambda) = E(\Lambda^+) \underset{\substack{\cup \\ \text{nm-disjoint}}}{\cup} E(\Lambda^-) \cup \underline{\underline{\text{Ends}}}$$

End: AdS_n Tetrahedron
 ||
 what can see x^-
 and can be seen from x^+
 (generator of Alexander Topology)

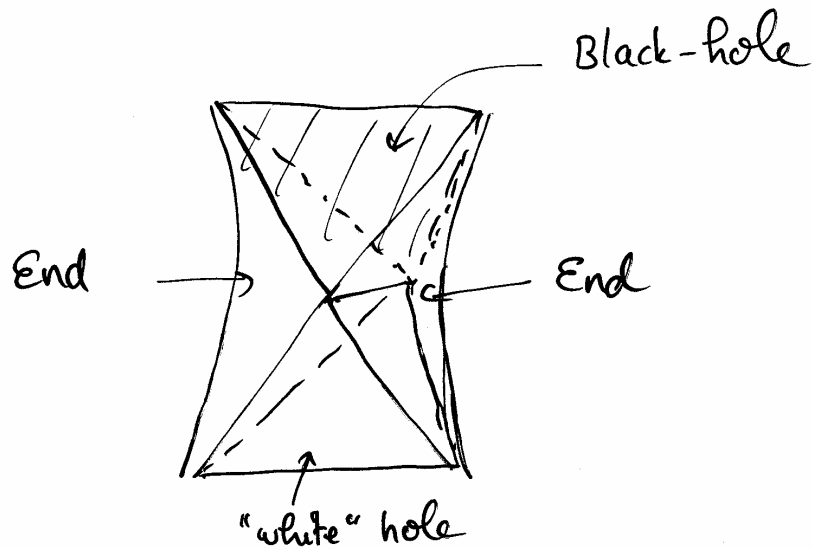


$E(\Lambda^+)$: Black-hole (Globally hyperbolic)

On end:

Kerr-like metric: $-(N^+)^2 dt^2 + (N^+)^{-2} dr^2 + r^2 (d\varphi + \frac{J}{2r^2} dt)^2$
 with $N^+ = \sqrt{\frac{1}{2} r^2 - M + \left(\frac{J}{2r}\right)^2}$

~~NON~~ STATIC Single BTZ black-hole
($J=0$)



$$\Gamma = \left(e^{\pi r_0 \Delta}, e^{\pi r_0 \Delta} \right)$$

$$\Delta = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

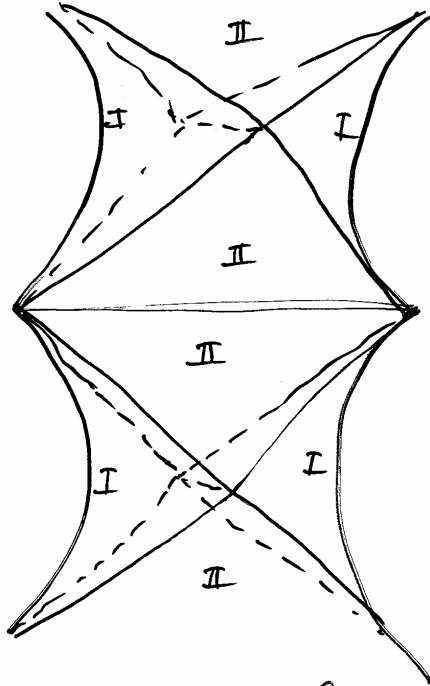
$$M = 2r_0^2$$

$$J = 0$$

Single BTZ black-hole :

Non Static

$$\mathbb{I} = \langle \gamma \rangle \text{ where } \mathbb{I}_0 = \left(\exp(\pi(r_+ - r_-)\Delta), \exp(\pi(r_+ + r_-)\Delta) \right)$$



$$M = r_+^2 + r_-^2$$

$$J = 2r_+ r_-$$

+ regions (III) : $D(\mathbb{I}) = \{ x \in \text{AdS} / \exists(z) \text{ spacelike} \}$
 where $\exists \gamma = \exp(z)$

Description similar to Kerr spacetime
 maximal

("extreme" case, ~ isolated)

In general:

Γ non abelian, admissible

$\forall \gamma \in \Gamma, \quad D(\gamma) = \{x / Z_\gamma(x) \text{ spacelike}\}$

$$D(\Gamma) = \bigcap_{\gamma \in \Gamma} D(\gamma)$$

THM: $D(\Gamma) = E(1, \rho)$

(Particular cases in Ammonberg, Bengtsson, Holst
"A spinning anti-de Sitter wormhole")

Last comments:

- Higher dimensions
- Wick/rescaling
rotation \rightarrow NON-COMPLETE HYPERBOLIC
MANIFOLDS (IF "MOMENTA IS
TOO BIG")
- Observables: $\text{Vol}(C(\text{an} \Gamma))$, etc...